# Markscheme 

# November 2018 

Physics

Higher level

## Paper 3

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## Section A

| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 1. | a | $m^{\frac{3}{2}} \checkmark$ | Accept other power of tens multiples of $m^{\frac{3}{2}}$, eg: $\mathrm{cm}^{\frac{3}{2}}$. | 1 |
| 1. | b | measured uncertainties «for one oscillation and for 20 oscillations» are the same/similar/OWTTE <br> OR <br> $\%$ uncertainty is less for 20 oscillations than for one $\checkmark$ <br> dividing «by 20» / finding mean reduces the random error $\checkmark$ |  | 2 |

(continued...)
(Question 1 continued)

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 1. | c | i | Straight line touching at least 3 points drawn across the range $\checkmark$ | It is not required to extend the line to pass through the origin. | 1 |
| 1. | c | ii | theory predicts proportional relation $« T \propto \frac{1}{d}$, slope $=T d=\frac{c}{\sqrt{g}}=$ constant » $\checkmark$ the graph is «straight» line through the origin $\checkmark$ |  | 2 |

(continued...)
(Question 1 continued)


| Question |  | Answers | Total |  |  |
| :--- | :--- | :--- | :--- | :--- | :--- |
| 2. | a |  | to provide a constant heating rate / power <br> OR <br> to have $m$ proportional to $t \checkmark$ |  |  |
| 2. | b |  | due to heat losses «VIt is larger than heat into liquid»$L_{v}$ calculated will be larger $\checkmark$ <br> 2. <br> c | heat losses will be similar / the same for both experiments <br> OR <br> heat loss presents systematic error $\checkmark$ <br> taking the difference cancels/eliminates the effect of these <br> losses <br> OR <br> use a graph to eliminate the effect $\checkmark$ | $\mathbf{2}$ |

## Section B

## Option A - Relativity

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 3. | a |  | a set of rulers and clocks / set of coordinates to record the position and time of events $\checkmark$ |  | 1 |
| 3. | b | i | ALTERNATIVE 1: <br> the time in frame $S^{\prime}$ is $t^{\prime}=\frac{L}{C} \checkmark$ but time is absolute in Galilean relativity so is the same in $S \checkmark$ <br> ALTERNATIVE 2: <br> In frame $S$, light rays travel at $c+v \checkmark$ <br> so $t=\frac{L}{(c+v)-v}=\frac{L}{c} \checkmark$ | In Alternative 1, they must refer to $S^{\prime}$ | 2 |
| 3. | b | ii | $x=x^{\prime}+v t \text { and } x^{\prime}=L \checkmark$ <br> «substitution to get answer» |  | 1 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 4. | a |  | $\begin{aligned} & \frac{0.82 c+0.40 c}{1+\frac{0.82 c \times 0.40 c}{c^{2}}} \\ & 0.92 c \checkmark \end{aligned}$ |  | 2 |
| 4. | b | i | $\begin{aligned} & \Delta t^{\prime}=\frac{120}{0.40 c} \checkmark \\ & \Delta t^{\prime}=1.0 \times 10^{-6} « \mathrm{~s} » \end{aligned}$ |  | 2 |
| 4. | b | ii | $\begin{aligned} & \gamma=« \frac{1}{\sqrt{1-0.82^{2}}}=» 1.747 \checkmark \\ & \Delta t=« \gamma\left(\Delta t^{\prime}+\frac{v \Delta x^{\prime}}{c^{2}}\right) »=1.747 \times\left(1.0 \times 10^{-6}+\frac{0.82 c \times 120}{c^{2}}\right) \end{aligned}$ <br> OR $\begin{aligned} & \Delta t=\frac{120}{1.747 \times(0.92-0.82) c} \\ & 2.3 \times 10^{-6} « \mathrm{~s} » \end{aligned}$ |  | 3 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 5. | a | i | $\begin{aligned} & \gamma=« \frac{1}{\sqrt{1-0.745^{2}}}=» 1.499 \checkmark \\ & x^{\prime}=« \gamma(x-v t)=» 1.499 \times(1.0-0) \checkmark \\ & « x^{\prime}=1.5 \mathrm{~m} » \end{aligned}$ |  | 2 |
| 5. | a | ii | $\begin{aligned} & t^{\prime}=« \gamma\left(t-\frac{v x}{c^{2}}\right)=» 1.499 \times\left(0-\frac{0.745 c \times 1}{c^{2}}\right) «=-\frac{1.11}{c} » \\ & « c t^{\prime}=-1.1 \mathrm{~m} » \end{aligned}$ <br> OR using spacetime interval $0-1^{2}=\left(c t^{\prime}\right)^{2}-1.5^{2} \Rightarrow « c t^{\prime}=-1.11 » \checkmark$ |  | 1 |

(continued...)
(Question 5 continued)

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 5. | b | i | line through event E parallel to $c t^{\prime}$ axis meeting $x^{\prime}$ axis and labelled $P$ J |  | 1 |
| 5. | b | ii | point on $x^{\prime}$ axis about $\frac{2}{3}$ of the way to $P$ labelled $Q \checkmark$ |  | 1 |

(continued...)
(Question 5 continued)


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 6. | a |  | $p c=\sqrt{E^{2}-\left(m c^{2}\right)^{2}}=\sqrt{1.50^{2}-0.511^{2}}$ « $=1.410 \mathrm{MeV}$ » |  | 1 |
| 6. | b | i | first equation is due to momentum conservation $\checkmark$ second equation is due to total energy conservation |  | 2 |
| 6. | b | ii | adding $2 p_{1}=3.42 \mathrm{MeV} \mathrm{c}^{-1} \Rightarrow p_{1}=1.71 \mathrm{MeV} \mathrm{c}^{-1} \checkmark$ $p_{2}=0.30 \mathrm{MeV} \mathrm{c}^{-1} \checkmark$ |  | 2 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 7. | a | i | the distance from the black hole at which the escape speed is the speed of light $\checkmark$ |  | 1 |
| 7. | a | ii | $R_{\mathrm{S}}=« \frac{2 G M}{c^{2}}=\frac{2 \times 6.67 \times 10^{-11} \times 4.0 \times 10^{36}}{9.0 \times 10^{16}}=» 5.9 \times 10^{9} « \mathrm{~m} » \downarrow$ |  | 1 |
| 7. | b |  | $2=\frac{1}{\sqrt{1-\frac{5.9 \times 10^{9}}{r}}}$ <br> rearranged to give $r$ <br> OR $\begin{aligned} & r=1.33 \times 5.9 \times 10^{9} \text { «m» } \\ & r=7.9 \times 10^{9} « \mathrm{~m} » \end{aligned}$ |  | 3 |

## Option B — Engineering physics

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 8. | a |  | taking torques about the pivot $R \times 4.00=36.0 \times 2.5 \checkmark$ $R=22.5 \text { «N» }$ |  | 2 |
| 8. | b | i | $\begin{aligned} & 36.0 \times 2.50=30.6 \times \alpha \checkmark \\ & \alpha=2.94 \text { «rad s}^{-2} » \end{aligned}$ |  | 2 |
| 8. | b | ii | the equation can be applied only when the angular acceleration is constant any reasonable argument that explains torque is not constant, giving non constant acceleration $\checkmark$ | eg weight is no longer perpendicular to the rod | 2 |
| 8. | c | i | «from conservation of energy» Change in GPE = Change in rotational KE $\checkmark$ $\begin{aligned} & W \frac{L}{2}=\frac{1}{2} I \omega^{2} \\ & \omega=\sqrt{\frac{36.0 \times 5.00}{30.6}} \\ & « \omega=2.4254 \mathrm{rad} \mathrm{~s}^{-1} » \end{aligned}$ |  | 3 |
| 8. | c | ii | $L=30.6 \times 2.43=74.4$ «Js» $\downarrow$ |  | 1 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 9. | a | i | ALTERNATIVE 1: $\begin{aligned} & P_{c}=P_{B}=\frac{P_{A} V_{A}}{V_{B}} \checkmark \\ & =\frac{2.8 \times 10^{6} \times 1 \times 10^{-4}}{2.8 \times 10^{-4}} «=1.00 \times 10^{6} \mathrm{~Pa} » \end{aligned}$ <br> ALTERNATIVE 2 $\begin{aligned} & 2.80 \times 10^{6} \times 1.00^{\frac{5}{3}}=P_{\mathrm{c}} \times 1.85^{\frac{5}{3}} \\ & P_{\mathrm{c}}=2.80 \times 10^{6} \times \frac{1.00^{\frac{5}{3}}}{1.85^{\frac{5}{3}}} «=1.00 \times 10^{6} \mathrm{~Pa} \text { » } \end{aligned}$ |  | 2 |
| 9. | a | ii | ALTERNATIVE 1: <br> Since $T_{B}=T_{A}$ then $T_{C}=\frac{V_{C} T_{B}}{V_{B}} \checkmark$ $=\frac{1.85 \times 385}{2.8} \text { «=254.4K» }$ <br> ALTERNATIVE 2: $\begin{aligned} & \frac{2.80 \times 1.00}{385}=\frac{1.00 \times 1.85}{T_{c}} « \mathrm{~K} » \checkmark \\ & T_{\mathrm{c}}=385 \times \frac{1.00 \times 1.85}{2.80} \text { « }=254.4 \mathrm{~K} » \checkmark \end{aligned}$ |  | 2 |

(continued...)
(Question 9 continued)

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 9. | b |  | $\begin{aligned} & \text { work done }=« p \Delta V=1.00 \times 10^{6} \times\left(1.85 \times 10^{-4}-2.80 \times 10^{-4}\right)=»-95 « \mathrm{~J} » \\ & \text { change in internal energy }=« \frac{3}{2} p \Delta V=-\frac{3}{2} \times 95=»-142.5 « \mathrm{~J} » \end{aligned}$ $\begin{aligned} & Q=-95-142.5 \checkmark \\ & \text { «-238 J» } \end{aligned}$ | Allow positive values. | 3 |
| 9. | c | i | net work is $288-238=50$ « $\mathrm{J} » ~ \checkmark$ efficiency $=« \frac{288-238}{288}=» 0.17$ Ј |  | 2 |
| 9. | c | ii | along $B \rightarrow C \checkmark$ |  | 1 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 10. | a |  | ice displaces its own weight of water / OWTTE <br> $O R$ <br> melted ice volume equals original volume displaced / OWTTE $\checkmark$ <br> no change will take place $\checkmark$ |  | 2 |
| 10. | b | i | continuity equation says $v \times A_{1}=u \times A_{2} \checkmark$ «and» $A_{1}=4 A_{2} \checkmark$ «giving result» |  | 2 |
| 10. | b | ii | Bernoulli: $\begin{aligned} & « \frac{1}{2} \rho v^{2}+\rho g H+P_{\mathrm{atm}}=\frac{1}{2} \rho u^{2}+0+P_{\mathrm{atm}} » \text { gives } \frac{1}{2} \times 1000 \times \frac{u^{2}}{16}+1000 \times 9.8 \times 5.0=\frac{1}{2} \times 1000 \times u^{2} \\ & u=10.2 « \mathrm{~m} \mathrm{~s}^{-1} » \checkmark \end{aligned}$ | Accept solving directly via conservation of energy. | 2 |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 11. | a | because the mass and the driver are out of phase «by $\pi » \downarrow$ so upwards $\checkmark$ | Justification needed for MP2 | 2 |
| 11. | b | ALTERNATIVE 1: $\begin{aligned} & « Q=2 \pi \frac{A_{0}^{2}}{A_{0}^{2}-A_{1}^{2}} » \Rightarrow \frac{A_{1}^{2}}{A_{0}^{2}}=1-\frac{2 \pi}{Q} \\ & \frac{A_{1}}{A_{0}}=« \sqrt{1-\frac{2 \pi}{22}}=» A_{1}=8.5 « \mathrm{~cm} » \end{aligned}$ <br> ALTERNATIVE 2: <br> driver amplitude is constant $\checkmark$ so mass amplitude is unchanged at $10 \mathrm{~cm} \checkmark$ |  | 2 |

## Option C - Imaging

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 12. | a |  | each incident ray shown splitting into two $\checkmark$ <br> each pair symmetrically intersecting each other on principal axis $\downarrow$ <br> for red, intersection further to the right $\checkmark$ |  <br> For MP3, at least one of the rays must be labelled. | 3 |
| 12. | b | i | rays diverge after passing through lens <br> OR <br> the extension of the rays will intersect the principal axis on the side of incident rays/as if they were coming from the focal point/points in the left side/OWTTE $\checkmark$ |  | 1 |
| 12. | b | ii | by placing a diverging lens next to the converging lens OR <br> make an achromatic doublet $\checkmark$ | Further details are not required. | 1 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 13. | a |  | proper construction lines $\downarrow$ <br> image at intersection of proper construction lines $\checkmark$ |  | 2 |

(Question 13 continued)


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 14. | a | i | $\begin{aligned} & « \sin \theta_{\mathrm{c}}=\frac{n_{1}}{n_{2}} » n_{1}=1.52 \times \sin 84.0^{\circ} \checkmark \\ & n_{1}=1.51 \checkmark \end{aligned}$ |  | 2 |
| 14. | a | ii | to have a critical angle close to $90^{\circ} \checkmark$ so only rays parallel to the axis are transmitted $\checkmark$ to reduce waveguide/modal dispersion $\checkmark$ | Do not accept "so that most rays are reflected". | 1 max |
| 14. | b | i | long path is $\frac{12 \times 10^{3}}{\sin 84^{\circ}} \checkmark$ $=12066 \text { « m» }$ <br> «so 66 m longer» |  | 2 |
| 14. | b | ii | speed of light in core is $\frac{3.0 \times 10^{8}}{1.52}=1.97 \times 10^{8} « \mathrm{~m} \mathrm{~s}^{-1} » \checkmark$ time delay is $\frac{66}{1.97 \times 10^{8}}=3.35 \times 10^{-7}$ «s» |  | 2 |
| 14. | b | iii | no, period of signal is $1 \times 10^{-8}$ «s » which is smaller than the time delay/OWTTE $\checkmark$ |  | 1 |


| Question |  | Answers | Notes | Total |
| :--- | :--- | :--- | :--- | :--- |
| 15. | a | protons spin direction changes <br> OR <br> proton energy state changes $\checkmark$ |  |  |
| 15. | b | Relaxation time «of signal/proton spin» $\checkmark$ <br> Location/time delay of the emitted RF signal $\checkmark$ |  |  |
| 15. | c |  | Relaxation time gives information on tissue type/density/health/OWTTE $\checkmark$ <br> Location information provides 3D image/OWTTE $\checkmark$ |  |


| Question |  | Answers | Notes | Total |  |
| :--- | :--- | :--- | :--- | :--- | :---: |
| 16. | a | $I_{0} e^{-23 \times 0.041} \checkmark$ <br> $=0.39 I_{0} \checkmark$ | $\mathbf{2}$ |  |  |
| 16. | b |  | $R=«\left(\frac{6.3 \times 10^{6}-1.7 \times 10^{6}}{6.3 \times 10^{6}+1.7 \times 10^{6}}\right)^{2}=» 0.33 \checkmark$ <br> so reflected intensity is $0.33 \times 0.39 I_{0}=0.13 I_{0} \checkmark$ |  |  |
| 16. | c | $0.13 I_{0} \times 0.39=0.05 I_{0} \checkmark$ | $\mathbf{2}$ |  |  |

## Option D - Astrophysics

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 17. | a |  | In cluster, stars are gravitationally bound $O R$ constellation not $\checkmark$ In cluster, stars are the same/similar age $O R$ in constellation not $\checkmark$ Stars in cluster are close in space/the same distance OR in constellation not $\checkmark$ <br> Cluster stars appear closer in night sky than constellation $\checkmark$ <br> Clusters originate from same gas cloud OR constellation does not $\checkmark$ |  | 2 max |
| 17. | b | i | $d=275$ «рс» $\checkmark$ |  | 1 |
| 17. | b | ii | because of the difficulty of measuring very small angles $\checkmark$ |  | 1 |
| 17. | C |  | mass of gas cloud > Jeans mass $\checkmark$ <br> «magnitude of» gravitational potential energy $>E_{k}$ of particles $\checkmark$ cloud collapses/coalesces «to form a protostar» $\downarrow$ |  | 2 max |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 18. | a | i | $\lambda=« \frac{2.9 \times 10^{-3}}{4600}=» 630 « n m » \downarrow$ |  | 1 |
| 18. | a | ii | black body curve shape $\downarrow$ peaked at a value from range 600 to $660 \mathrm{~nm} \sqrt{ }$ |  | 2 |
| 18. | a | iii | $\begin{aligned} & \frac{L}{L_{\odot}}=\left(\frac{0.73 R_{\odot}}{R_{\odot}}\right)^{2} \times\left(\frac{4600}{5800}\right)^{4} \checkmark \\ & L=0.211 L_{\odot} \checkmark \end{aligned}$ |  | 2 |
| 18. | b |  | $M=« 0.21^{\frac{1}{3.5}} M_{\odot}=» 0.640 M_{\odot} \checkmark$ | Accept reverse argument $0.644^{3.5}=0.21$ | 1 |
| 18. | C |  | $\frac{T_{E}}{T_{\odot}}=« \frac{\frac{M_{E}}{L_{E}}}{\frac{M_{\odot}}{L_{\odot}}}=\frac{0.64}{0.21}=» 3.0 \checkmark$ <br> $T \approx 27$ billion years $\checkmark$ |  | 2 |
| 18. | d |  | red giant $\checkmark$ <br> planetary nebula $\checkmark$ <br> white dwarf $\sqrt{ }$ | do NOT accept supernova, red supergiant, neutron star or black hole as stages | 3 |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 19. | a | measured redshift «z» of star $\checkmark$ use of Doppler formula $O R z \sim v / c$ $O R v=\frac{c \Delta \lambda}{\lambda}$ to find $v \checkmark$ | OWTTE | 2 |
| 19. | b | use of gradient or any point on the line to obtain any expression for either $H=\frac{v}{d}$ or $t=\frac{d}{v} \checkmark$ correct conversion of $d$ to m and v to $\mathrm{m} / \mathrm{s} \checkmark$ $=4.6 \times 10^{17} \text { «s » }$ |  | 3 |


| Question |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: |
| 20. | a | energy filling all space $\checkmark$ resulting in a repulsive force/force opposing gravity $\checkmark$ accounts for the accelerating universe $\checkmark$ makes up about $70 \%$ of «the energy» of universe $\sqrt{ }$ |  | 2 max |
| 20. | b | black hole $\checkmark$ brown dwarf $\downarrow$ massive compact halo object /MACHO $\sqrt{ }$ neutrinos $\checkmark$ weakly interacting massive particle /WIMP $\checkmark$ |  | 2 max |


| Question |  | Answers | Notes | Total |
| :--- | :--- | :--- | :--- | :--- | :---: |
| 21. | a | «wavelength of light/CBR» $\lambda \propto R \checkmark$ <br> reference to Wien's law showing that $\lambda \propto \frac{1}{T} \checkmark$ <br> combine to get result $\checkmark$ | OWTTE |  |
| 21. | b | $\frac{R_{\text {past }}}{R_{\text {now }}}=\frac{3}{300}=0.01 \checkmark$ 3 |  |  |

